

QUANTUM POSITIONING SYSTEM

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A quantum positioning system (QPS) is proposed that can provide a user with all four of his space-time coordinates. The user must carry a corner cube reflector, a good clock, and have a two-way classical channel of communication with the origin of the reference frame. Four pairs of entangled photons (biphotons) are sent through four interferometers: three interferometers are used to determine the user's spatial position, and an additional interferometer is used to synchronize the user's clock to coordinate time in the reference frame. The spatial positioning part of the QPS is similar to a classical time-of-arrival (TOA) system; however, a classical TOA system (such as GPS) must have synchronized clocks that keep coordinate time and, therefore, the clocks must have long-term stability, whereas in the QPS only a photon coincidence counter is needed and the clocks need only have short-term stability. Several scenarios are considered for a QPS: one is a terrestrial system and another is a space-based system composed of low-Earth orbit (LEO) satellites. Calculations indicate that for a space-based system, neglecting atmospheric effects, a position accuracy below the 1 cm level is possible for much of the region near the Earth. The QPS may be used as a primary system to define a global four-dimensional reference frame.

I. INTRODUCTION

During the past several years, the Global Positioning System (GPS) has practically become a household word. The GPS is a U.S. Department of Defense satellite system that is used by both the military and civilians for navigation and time dissemination [1-3]. Automobile, ship, aircraft, and spacecraft use the GPS for navigation. Telephone and computer network systems that require precise time use the GPS for time synchronization. The GPS is a complex system consisting of approximately 24 satellites orbiting the Earth in circular orbits at approximately 4.25 Earth radii. The GPS is designed so that signals travel line-of-site from satellite to user, and from any place on Earth at least four satellites are in view. If a user receives four GPS satellite signals simultaneously from four satellites, $s = 1, 2, 3, 4$, and the satellites' space-time coordinates (t_s, \mathbf{r}_s) at time of transmission are known, the user can solve for his unknown space-time coordinates, (t_o, \mathbf{r}_o) , by solving the four equations [4, 5].

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$$|\mathbf{r}_o - \mathbf{r}_s|^2 - c^2(t_o - t_s)^2 = 0, \quad s = 1, \dots, 4 \quad (1)$$

In Eq. (1) we assume that the signals travel on four light cones that are centered at the reception event and we have ignored atmospheric delays. The actual signals that the GPS satellites transmit are continuous-wave circularly polarized radio-frequency signals on two carrier frequencies in the L-band centered about: $L_1 \approx 1575.42$ MHz and $L_2 \approx 1227.6$ MHz. The carrier-frequency signals are modulated by a pseudorandom noise (PRN) code. A GPS receiver makes a phase-difference measurement, called a *pseudorange measurement*, which is the phase difference between the PRN code received from the satellite and an identical copy of the PRN code that is replicated inside the GPS receiver; see [5] for details of the GPS pseudorange measurement process. The pseudorange measurement is essentially made by performing a correlation of the code bits in the PRN code received from the satellite with an identical copy of the same PRN code replicated inside the GPS receiver.

Recently, there have been several proposals for synchronizing clocks by making use of entangled quantum systems [6-12]. The question naturally arises whether entangled quantum systems can be exploited to determine all four space-time coordinates of a user, rather than just time.

In this article, I describe a quantum positioning system (QPS) that is in principle capable of providing a user with all four of his space-time coordinates. This QPS is the quantum analog of the classical GPS described above. The QPS is based on entangled photon pairs (biphotons) and second order correlations within each pair. The two-photon coincidence counting rate is the basic measured quantity. In order to determine his four space-time coordinates, a user of the QPS must carry a corner cube reflector, a good clock, and have a two-way classical channel for communication with the origin of the reference frame.

II. INTERFEROMETRIC QUANTUM POSITIONING SYSTEM

For simplicity of discussion, I assume that space-time is flat with Minkowski [5] coordinates (t, x, y, z) and that the user of the QPS is stationary. The complete QPS consists of four biphoton sources (entangled photon pairs), four beam splitters and four two-photon coincidence counting Hong-Ou-Mandel (HOM) interferometers; see Figure 1. Three of the interferometers are used together to solve for the user's position and one interferometer is used to solve for the user's time, in a particular reference frame.

Six spatial points, \mathbf{R}_i , where $\mathbf{R}_i = (x_i, y_i, z_i)$, for $i = 1, 2, 3, \dots, 6$, define the spatial part of the reference frame at constant coordinate time t . The six points \mathbf{R}_i define three independent baselines in pairs, $(\mathbf{R}_1, \mathbf{R}_2)$, $(\mathbf{R}_3, \mathbf{R}_4)$, and $(\mathbf{R}_5, \mathbf{R}_6)$. The points \mathbf{R}_i are assumed to be accurately surveyed, so their coordinates are precisely known. Determination of a user's spatial coordinates, $\mathbf{r}_o = (x_o, y_o, z_o)$, is done with respect to this reference frame. A stationary clock in this reference frame, say at the origin of coordinates, $(x, y, z) = (0, 0, 0)$, provides a measure of coordinate time t in this four-dimensional system of coordinates. We neglect all gravitational effects [4] so that the

user's clock, which keeps proper time τ , runs at the same rate as coordinate time t in the reference frame defined by the spatial points \mathbf{R}_i , so that $d\tau/dt = 1$. Synchronization of the user's clock to coordinate time [13] means that we have a method to compute the integration constant τ_o in $\tau - \tau_o = t$. In four-dimensional flat space-time, the world lines of the spatial points \mathbf{R}_i define a tube. Events that are simultaneous in this system of four-dimensional coordinates occur at constant values of coordinate time t , which is a hyperplane that cuts this tube.

There are several possible modes for the QPS. First, consider a user that must determine all four of his space-time coordinates (t_o, x_o, y_o, z_o) . In a previous paper [12], an algorithm has been given to synchronize the proper time τ on a user's clock to coordinate time, t , without prior knowledge of the geometric range from the reference clock (which keeps coordinate time t) to the user clock. We assume that this algorithm is employed here to synchronize the user's clock with coordinate time t . This algorithm requires a two-way classical channel of communication between the user and the reference frame origin, where coordinate time t is kept. The three spatial coordinates are determined separately as follows (refer to Figure 1).

Each baseline, such as the one connected by points, $(\mathbf{R}_1, \mathbf{R}_2)$, contains an entangled photon pair (biphoton) source [14-18] located in the baseline. For convenience, we take the biphoton source to be at the midpoint of the baseline at point E_1 at position \mathbf{r}_1 . Additionally, each baseline contains a 50:50 beam splitter and two photon detectors; see Figure 2. For simplicity of discussion, we assume that the biphoton source is essentially collocated with the beam splitter and two photon detectors at point E_1 ; see Figure 2. Along the baseline there is a controllable optical delay at point D_1 . The other two baselines contain the same equipment, as shown in Figure 2. The QPS works as follows.

Photon pairs are created at E_1 , are sent to positions \mathbf{R}_1 and \mathbf{R}_2 , and are redirected to the user at the unknown position \mathbf{r}_o . The two photon paths are similar, except that one path has a controllable optical delay D_1 . The optical delay is assumed to be calibrated so that we can accurately impose an arbitrary delay time [19]. Next, the entangled photons reflect from the user's corner cube reflector at \mathbf{r}_o , and return back through the same paths, through points \mathbf{R}_1 and \mathbf{R}_2 , and arrive at a HOM interferometer that is collocated at E_1 at position \mathbf{r}_1 ; see Figure 2. For convenience, we assume that the interferometer is collocated with the photon generation point \mathbf{r}_1 . Again, both photon return paths are similar, but one path has the optical delay D_1 . We have the following effective round-trip times for each photon path

$$t_L = \frac{2}{c} [|\mathbf{r}_o - \mathbf{R}_1| + |\mathbf{R}_1 - \mathbf{r}_1|] \quad (2)$$

$$t_R = \frac{2}{c} [|\mathbf{r}_o - \mathbf{R}_2| + |\mathbf{r}_1 - \mathbf{R}_2| + (n - 1) d]$$

where d is the geometric thickness of the optical delay D_1 perpendicular to the optical path and n is the effective index of refraction for the optical delay D_1 . The optical delay D_1 is now adjusted until a minimum is observed in the two-photon counting rate at E_1 [15]. When the minimum in photon coincidence counting rate is observed at interferometer E_1 , the effective travel times for each photon path are equal, $t_L = t_R$. The interferometer is balanced when the condition $t_L = t_R$ is satisfied, and a unique minimum is observed in the two-photon counting rate R_c . The accuracy with which this minimum can be observed depends on the bandwidth $\Delta\omega$ of the band-pass interference

filters used in front of the photon detectors.

We get an equation that relates the geometric path lengths to the measured optical delay time $\Delta t_1 = (n - 1)d/c$:

$$|\mathbf{r}_o - \mathbf{R}_1| + |\mathbf{R}_1 - \mathbf{r}_1| = |\mathbf{r}_o - \mathbf{R}_2| + |\mathbf{r}_1 - \mathbf{R}_2| + c\Delta t_1 \quad (3)$$

An analogous measurement process is carried out for the other two baselines. For simplicity, I assume that the points E_1 , E_2 , and E_3 are located at midpoints of their baselines. We then obtain the three equations

$$|\mathbf{r}_o - \mathbf{R}_1| = |\mathbf{r}_o - \mathbf{R}_2| + s_1 \quad (4)$$

$$|\mathbf{r}_o - \mathbf{R}_3| = |\mathbf{r}_o - \mathbf{R}_4| + s_2 \quad (5)$$

$$|\mathbf{r}_o - \mathbf{R}_5| = |\mathbf{r}_o - \mathbf{R}_6| + s_3 \quad (6)$$

where $s_i = c\Delta t_i$ for $i = 1, 2, 3$.

The two-photon coincidence counting rate is given by [20-22]

$$R_c = \alpha_1 \alpha_2 |\eta V|^2 |G(0)|^2 [1 - e^{-(\Delta\omega \Delta t_1)^2}] \quad (7)$$

where $|V|^2$ is the pump intensity in photons per second, α_1 and α_2 are the quantum efficiencies of detectors D_1 and D_2 , η is a dimensionless constant and $G(t)$ is the Fourier transform of the spectral function ϕ , which is the autocorrelation function of the down-converted light

$$G(t) = \int_0^\infty \phi\left(\frac{1}{2}\omega_0 + \omega, \frac{1}{2}\omega_0 - \omega\right) e^{-i\omega t} d\omega \quad (8)$$

The Eqs. (4)-(6) can be solved for the three unknown user spatial coordinates, $\mathbf{r}_o = (x_o, y_o, z_o)$, in terms of the three measured time delays, $\Delta t_1, \Delta t_2, \Delta t_3$, which balanced the three interferometers. The measured data consists of photon coincidence count rate vs. optical time delay lengths s_i , for $i = 1, 2, 3$. Clearly a search must be done of the data to locate the minimum that corresponds to equal time of travel along the interferometer arms. The computations can be done at points E_1, E_2 , and E_3 . This search to locate the minimum is the quantum analog of the correlation of the PRN code in a classical GPS receiver, which was described above. When the three interferometers at E_1, E_2 , and E_3 have been balanced simultaneously, a classical message is sent to the user giving him the values of his coordinates $\mathbf{r}_o = (x_o, y_o, z_o)$. Clearly, classical communication is needed between the points $\mathbf{r}_1, \mathbf{r}_2$, and \mathbf{r}_3 to establish that the interferometers are balanced at a given coordinate time t . We imagine that when each interferometer is balanced, a message is sent to the origin of coordinates. When three messages are simultaneously received at the origin of coordinates (saying that the three interferometers are balanced), Eqs. (4)-(6) are solved for the user's coordinates $\mathbf{r}_o = (x_o, y_o, z_o)$ and the user's coordinates are sent to the user through a classical channel of communication.

In the QPS that we describe, there is an apparent asymmetry in the determination of a user's spatial coordinates, $\mathbf{r}_o = (x_o, y_o, z_o)$, and in the determination of the user's time. In my view, this asymmetry is a reflection of the asymmetric way that space and time enter in the theory of the quantized electromagnetic field to give rise to photons as quanta of the field. As mentioned above, the time synchronization of the user's clock to coordinate time is done by a method previously described by Bahder and Golding [12]. Therefore in what follows, I discuss only the spatial part of the QPS.

With some modification of the above scheme, we may imagine that we could design a similar system based on first-order coherence for position determination [20, 21]. A single beam from a continuous-wave laser can be split and the beams sent on two different paths. However, in such a case, there would be an ambiguity that is associated with the wavelength of the light (interference fringes will be seen) that is unresolvable in principle. In contrast, in the quantum case (which relies on second-order coherence) the ambiguity is resolved because equal propagation times for two paths lead to quantum interference: equal travel times for two paths create a *unique* observable minimum in the two-photon coincidence counting rate R_c .

The measured quantities in the QPS are the optical path delays s_i . For a given measured value of optical delay, say s_1 , Eq.(4) specifies that the user's coordinates must lie on a hyperboloid surface with foci at \mathbf{R}_1 and \mathbf{R}_2 , i.e., a hyperbola of revolution that is symmetric about the baseline defined by \mathbf{R}_1 and \mathbf{R}_2 . The user's position, \mathbf{r}_o , is then given by the intersection of three hyperbolae given by Eqs. (4)-(6). Each Eq. (4)-(6) is just the equation for a baseline in a classical time of arrival (TOA) system that records arrival times of classical light pulses (or distinct intensity edges) at two spatial reception points \mathbf{R}_i . In the case of a classical TOA system, pulse arrival time at *four* locations is needed to determine all four space-time coordinates. In that case, four time difference of arrival (TDOA) equations can be formed from four points, and each point is used multiply to (effectively) form the baselines. (Taking TDOAs results in a system of three equations where the emission event time has cancelled out.) In the quantum case, since correlations between photon pairs are used, we must use three baselines defined by six points \mathbf{R}_i , plus an additional interferometer for the determination of the user's time. As we will see below, the QPS is an interferometric system.

More fundamentally, and more significant for applications, is that in the classical case of a TOA system, we must have good clocks that are synchronized to coordinate time so that accurate pulse arrival times at the four \mathbf{R}_i reception points can be recorded. A good clock that keeps coordinate time is often a difficult requirement to meet in practice [23]. In contrast, in the quantum case two-photon coincidence counts at detectors D_1 and D_2 are measured and only a good "flywheel" clock is needed (i.e., a clock having a good short-term stability) to measure photon coincidence count rates while the optical time delay is adjusted, to locate the minimum in R_c .

While the QPS that I describe is similar to the classical GPS, there are significant differences. In the case of the GPS, if a user does not know his coordinate time, then he must observe signals from four satellites, even if he only wants to determine his spatial coordinates. In the case of the QPS, a user can determine his spatial coordinates without having to determine his coordinate time (by using three HOM interferometers). A more fundamental distinction between QPS and GPS is

that the GPS is a one-way signal travel system, whereas the QPS is a two-way system. Furthermore, in the GPS classical information (consisting of emission event time at the satellite) is carried in the PRN code bit numbers broadcast by the GPS satellites, whereas in the QPS no information is carried by the photons. Therefore, it seems that there is no classical analog of the QPS that can be implemented with PRN codes (which are deterministic and periodic in time).

III. GEOMETRIC DILUTION OF PRECISION

In the case of the classical GPS, the geometrical positions of the GPS satellites determine the accuracy of the user's position. This effect is sometimes called the geometric dilution of precision (GDOP). We compute the positioning accuracy and the effect of GDOP for the QPS from Eqs. (4)-(6). These equations give an implicit relation $\mathbf{r}_o = \mathbf{r}_o(\mathbf{R}_1, \mathbf{R}_2, \mathbf{R}_3, \mathbf{R}_4, \mathbf{R}_5, \mathbf{R}_6, s_1, s_2, s_3)$ for the user position \mathbf{r}_o as a function of the three measured path delays, s_i , and the six baseline endpoints \mathbf{R}_i . If we knew the error in the measured path length delays, ds_1 , ds_2 , and ds_3 , we could compute the error in the three components of the user's position vector, $d\mathbf{r}_o = (dx_o, dy_o, dz_o)$, from

$$d\mathbf{r}_o = \sum_{i=1}^3 \frac{\partial \mathbf{r}_o}{\partial s_i} ds_i \quad (9)$$

for constant \mathbf{R}_i . However, these errors are statistical in nature, so instead I compute the standard deviations σ_x , σ_y , and σ_z , of the user coordinates x_o , y_o , and z_o , as a function of the standard deviations σ_{s_1} , σ_{s_2} , and σ_{s_3} , of the measured optical time delays s_1 , s_2 , and s_3 . For constant \mathbf{R}_k for $k = 1, \dots, 6$, these standard deviations are related by [24]

$$\begin{aligned} \sigma_x^2 &= \left(\frac{\partial x_o}{\partial s_1} \right)^2 \sigma_{s_1}^2 + \left(\frac{\partial x_o}{\partial s_2} \right)^2 \sigma_{s_2}^2 + \left(\frac{\partial x_o}{\partial s_3} \right)^2 \sigma_{s_3}^2 \\ \sigma_y^2 &= \left(\frac{\partial y_o}{\partial s_1} \right)^2 \sigma_{s_1}^2 + \left(\frac{\partial y_o}{\partial s_2} \right)^2 \sigma_{s_2}^2 + \left(\frac{\partial y_o}{\partial s_3} \right)^2 \sigma_{s_3}^2 \\ \sigma_z^2 &= \left(\frac{\partial z_o}{\partial s_1} \right)^2 \sigma_{s_1}^2 + \left(\frac{\partial z_o}{\partial s_2} \right)^2 \sigma_{s_2}^2 + \left(\frac{\partial z_o}{\partial s_3} \right)^2 \sigma_{s_3}^2 \end{aligned} \quad (10)$$

where the partial derivatives are done at constant \mathbf{R}_k . The lengthy calculation to compute the partial derivatives in Eq. (10) is done analytically using *Mathematica*. For simplicity, I assume that the error distributions of the s_i are Gaussian and that the three standard deviations are equal, $\sigma_{s_1} = \sigma_{s_2} = \sigma_{s_3} \equiv \sigma_s$. For a spherically symmetric probability distribution of three-dimensional positions $\mathbf{r}_o = (x_o, y_o, z_o)$, the spherical error probable (SEP), which is the radius R within which 50% of the points lie, is related [25] to the standard deviations $\sigma_x = \sigma_y = \sigma_z \equiv \sigma$ by $R \cong 1.538\sigma$. In our case, the probability distribution of \mathbf{r}_o is not necessarily spherical. To approximate the SEP error metric, we compute a weighted approximation to the SEP metric by defining $R_{xyz} \cong 1.538 \frac{1}{\sqrt{3}} (\sigma_x^2 + \sigma_y^2 + \sigma_z^2)^{1/2}$. When the error distribution for \mathbf{r}_o is spherically symmetrical, the error metrics are equal: $R_{xyz} = R$. I consider the effect of GDOP for two cases, one in which the interferometer baselines are near each other, and the other case where the baselines are well separated, which is the case with classical GPS or a classical TOA system.

A. Geodetic Positioning System

First, consider a case where the three baselines coincide with the three Cartesian coordinate axes of a reference frame; see Figure 3. Such a case might occur when the baselines are on the Earth, and we want to determine the position of an object with respect to a topocentric coordinate system. For example, consider the center of the QPS at the origin of Cartesian coordinates and an object with a corner reflector at a range of 100m from the QPS, with coordinates $(x_o, y_o, z_o) = (100 \text{ m})(1, 1, 1)/\sqrt{3}$. Figure 4 shows a plot of contours of constant values of $1/R_{xyz}$ in the $x_o - y_o$ plane at $z_o = 100/\sqrt{3} \text{ m}$, for the interferometer arm (half) length $a = 2 \text{ m}$ and error (standard deviation) in optical path $\sigma_s = 1.0 \times 10^{-6} \text{ m}$. In the contour plot, the position error is $R_{xyz} = 8.3 \text{ cm}$ for $(x_o, y_o, z_o) = (100 \text{ m})(1, 1, 1)/\sqrt{3}$, whereas for $(x_o, y_o, z_o) = (30 \text{ m}, 30 \text{ m}, 100/\sqrt{3})$ the error $R_{xyz} = 3.9 \text{ cm}$, which corresponds to the upper right high-accuracy (light-shaded) region in Figure 4. On the z -axis at $x_o = y_o = 0$ and $z_o = 100 \text{ m}$ the error R_{xyz} is essentially infinite. Figure 5 shows a plot of the error metric R_{xyz} vs. x_o , for $y_o = 30 \text{ m}$ and $z_o = 100 \text{ m}/\sqrt{3}$, which corresponds to a line in Figure 4 with relatively small error R_{xyz} . In the high-accuracy light-shaded region of Figure 4, for $x_o = y_o = 30 \text{ m}$ and $z_o = 100 \text{ m}/\sqrt{3}$, the dependence of the error R_{xyz} on the baseline length $2a$ is plotted in Figure 6, also using $\sigma_s = 1.0 \times 10^{-6} \text{ m}$. For a four-meter baseline, $2a = 4 \text{ m}$, the error is just under 5 cm.

Note that the position error R_{xyz} depends linearly on σ_s , which is the standard deviation (error) in measurement of the optical path delay needed to obtain the minimum in two-photon coincidence counts R_c . The width of this minimum depends on the interference filters in front of the photon coincidence counting detectors as well as the pump laser bandwidth [15, 26]. Depending on the experimental design, this minimum may be measured to better than $\sigma_s = 1.0 \times 10^{-6} \text{ m}$, which was used in these plots, and hence accuracies may be better than plotted.

Finally, I note that the error function R_{xyz} has a very complex dependence on user coordinates (x_o, y_o, z_o) , and as stated earlier, the error function R_{xyz} also depends critically on the way the baselines are distributed, i.e., it depends on the six points \mathbf{R}_k for $k = 1, \dots, 6$, which define the baseline endpoints. In the next example, I consider a situation where the baselines do not intersect, and thereby the error R_{xyz} is considerably smaller than for the case considered above, even though the distances are larger.

B. Satellite-Based QPS

Now assume that each point of a baseline, \mathbf{R}_i , is associated with a different satellite, and that the spatial interferometer legs are formed from pairs of points $(\mathbf{R}_1, \mathbf{R}_2)$, $(\mathbf{R}_3, \mathbf{R}_4)$, and $(\mathbf{R}_5, \mathbf{R}_6)$; see Figure 7. Specifically, I assume that the points \mathbf{R}_k , are on low-Earth orbit (LEO) satellites. It may seem optimistic that a QPS is feasible with such large baselines because single photons must be propagated over these baselines and then reliably detected. However, recently single photons have been propagated through the atmosphere and detected over 10 km distance in daylight [27], and another study concludes that there are no obstacles to create a single-photon quantum key distribution system between ground and low-Earth orbiting satellites [28]. Therefore, a LEO-satellite QPS may be possible.

As an example of the positioning errors in a QPS made from LEO satellites, I take the baseline endpoints to be: $\mathbf{R}_1 = (a, -b/2, 0)$, $\mathbf{R}_2 = (a, b/2, 0)$, $\mathbf{R}_3 = (b/2, a, 0)$, $\mathbf{R}_4 = (-b/2, a, 0)$, $\mathbf{R}_5 = (-b/(2\sqrt{2}), -b/(2\sqrt{2}), a)$, and $\mathbf{R}_6 = (b/(2\sqrt{2}), b/(2\sqrt{2}), a)$. A plot of this configuration is shown in Figure 7. A contour plot of the reciprocal error function, $1/R_{xyz}$, is shown in the $x_o - y_o$ plane for $z_o = R_e/\sqrt{3}$, where $R_e = 6378$ km is the Earth's radius, see Figure 8. As an example, in the calculations below I take the semi-major axis of the LEO satellites to be $a = 7360$ km and the baseline between pairs of satellites as $b = 20$ km. The standard deviation (error) in the measured optical delay is taken to be $\sigma_s = 1.0 \mu\text{m}$. For a user on the surface of the Earth with coordinates $(x_o, y_o, z_o) = (1, 1, 1)R_e/\sqrt{3}$ the error is $R_{xyz} = 0.10$ cm. For these same parameters, Figure 9 shows a plot of the position error R_{xyz} vs. x_o for $y_o = z_o = R_e/\sqrt{3}$. Note that over a large range of x_o -values the error remains below 1 cm. Finally, Figure 10 shows a plot of the position error in the radial direction: $R_{xyz}(r_o/\sqrt{3}, r_o/\sqrt{3}, r_o/\sqrt{3})$ vs. r_o for the same parameters. On a radial line in the (1,1,1) direction, the error remains below 1 cm up to $r_o = 11680$ km. However, near 1300 km the error rises steeply. This is an example of the complex dependence of R_{xyz} on user position, which was mentioned earlier.

Clearly, the geometric positioning and layout of the baselines significantly affects the accuracy of a user's position. Note that the terrestrial QPS (discussed in the previous section) had a ratio of baseline length to user position $a/r_o = 0.02$, whereas this LEO satellite QPS has $b/a = 0.003$. By comparing the baseline layout for the terrestrial QPS and this LEO satellite QPS, it is clear that the positioning accuracy is sensitive to the separation and layout of the baselines, but not so sensitive to the baseline lengths. Other calculations (not shown) support this conclusion.

The above calculations for a satellite-based QPS are only meant as an example to illustrate the magnitude of errors in position that may be achievable. A significant amount of engineering calculations must be performed to design a realist satellite-based QPS. Furthermore, real satellites are moving and engineering similar to that used in the classical GPS would have to be done, e.g., using Kalman filtering techniques. Obviously, bright sources of entangled photons (biphotons) are needed. The calculations above suggest that if properly engineered, a satellite-based QPS may achieve position accuracy of objects near the Earth's surface below 1 cm. In these calculations, I have ignored the time delays introduced by the atmosphere. However, corrections can be made for atmospheric effects using multiple colors of photons similarly to what is done with the GPS. Perhaps one advantage of the quantum system as compared to the classical GPS is that entangled photons exhibit group velocity dispersion cancellation, which may be an important factor for future engineering and design of a QPS [29-32].

IV. ALTERNATIVE SCENARIOS

A. Position-Only Determination

A QPS can be designed to work in several modes, depending on the needs of the user and the required scenario. In the above discussion, we have described the case where a user of the QPS wants to determine both his spatial and time coordinates, (t_o, x_o, y_o, z_o) . A second alternative is that a user may only need to obtain their spatial coordinates, and he may not need the correct

time. In this latter case, the time synchronization portion of the system is not needed, and the user may find their position coordinates (x_o, y_o, z_o) by having only a corner reflector and a one-way (reception only) classical channel of communication with the reference frame origin, where the simultaneity of the three two-photon coincidence counting rate minima is established.

Another mode of operation of a QPS is where we want to determine the position of an object with a corner cube reflector, such as a geostationary satellite. In such a case, information on the position of the satellite, $\mathbf{r}_o = (x_o, y_o, z_o)$, is only needed on the ground. The satellite's position coordinates can be determined on the ground using a QPS, and only a corner cube reflector is needed on the satellite, but no communication channel to the satellite and no on-board clock is needed.

B. User Carries QPS Receiver

The scenarios that we have described above are ones where the measurements (adjusting the optical delays) and the calculations (to compute \mathbf{r}_o) are done near the origin of the reference frame. In a classical GPS receiver, the computations (correlations of PRN codes to at least four satellites) are done locally in the user's GPS receiver that the user carries with him. The QPS analog of this classical GPS scenario is a setup where the biphotons are generated at points E_1 , E_2 , and E_3 , but the user carries with him the 50:50 beam splitters and photon detectors. In this scenario, the user controls (and carries with him) the optical delays (see Figure 1), and he locally measures the optical delays s_1 , s_2 , and s_3 . The user must receive a classical message consisting of the coordinates of baseline endpoints, \mathbf{R}_i , $i = 1, \dots, 6$, and then he must solve the Eqs. (4)-(6) such a case, there are no clocks on-board the broadcasting satellites (located at positions \mathbf{R}_i), however, the user must carry a clock with short term stability to determine rate of photon coincidence counts from each of the three baselines (associated with spatial positioning) and also he must do coincidence counting for time determination (if time is needed). For the three spatial baselines, optical propagation is then one-way (using the satellite positions as a primary reference system, see below) from satellites to QPS user receiver. For time synchronization, however, as mentioned previously, the optical propagation must be two-way (when using the method of Bahder and Golding). In essence, for each of the four channels, the QPS receiver consists of a beam splitter, two single-photon detectors, and a controllable optical delay. All four space-time coordinates can be obtained by a user in this way. One clock in the reference frame must have long-term stability to define coordinate time, and another clock in the QPS user receiver can have short-term stability. *Note that the satellites do not need to carry clocks, because their positions can be used to define the primary system of coordinates.* This type of QPS is a close analog of the classical GPS.

V. QPS SPACE-TIME COORDINATES

The satellites at baseline points \mathbf{R}_i can be taken to define the *primary* system of reference, even though the points \mathbf{R}_i change with time. The quantities measured by a user of such a QPS are then (s_0, s_1, s_2, s_3) , where s_0 is the optical time delay (in the HOM interferometer) that will provide the user with coordinate time in this coordinate system (using the Bahder and Golding method), and (s_1, s_2, s_3) are the three optical delays in the three interferometers for position determination. The quantities (s_0, s_1, s_2, s_3) are then to be regarded as generalized four-dimensional space-time

coordinates [33], s_0 is a time-like coordinate and (s_1, s_2, s_3) are space-like coordinates. Within the context of general relativity, such coordinates are as good as any other coordinates, and they enter into the metric $c^2 d\tau^2 = g_{ij} ds_i ds_j$ of the flat space-time assumed in this work. Of course, a transformation from the QPS space-time coordinates, (s_0, s_1, s_2, s_3) , to an Earth-centered inertial (ECI) system of coordinates, say (t, x, y, z) , is of interest for astrodynamical applications. Such a transformation can be done approximately by conventional means of tracking the satellites (at baseline points \mathbf{R}_i).

It is interesting to remark that the QPS allows the direct measurement of four-dimensional space-time coordinates. Previously, it was believed that space-time coordinates were not measurable quantities [33-36]. Of course, the QPS coordinates (s_0, s_1, s_2, s_3) are real physical measurements, and it is well-known that real measurements are space-time invariants under generalized coordinate transformations [33].

VI. SUMMARY

I have presented a conceptual scheme for an interferometric quantum positioning system (QPS) based on second-order quantum coherence of entangled photon pairs (biphotons). A user's spatial coordinates are determined by locating three unique minima in three different two-photon counting rates, associated with three HOM interferometers built on independent baselines. The spatial portion of the QPS is similar to a classical TOA system; however, a classical TOA system requires synchronized clocks that keep coordinate time, which is often a difficult requirement to meet. In contrast, the QPS only requires a clock having a short-term stability to measure two-photon coincidence counting rates while the optical time delay is adjusted (to locate the minima in the two-photon coincidence counting rate R_c). Bright sources of entangled photons (biphotons) are needed.

Several different scenarios were considered for a QPS: one is a terrestrial system and the other is space-based. In both cases, I computed the accuracy of a user's position as a function of user position. The function that describes the errors in position has a complex spatial dependence. In the case of the terrestrial QPS, the position accuracy was relatively poor because the baselines were located near each other. This could be dramatically improved by moving apart the baselines.

As an example of a satellite-based QPS, I have proposed a LEO-satellite QPS. Neglecting atmospheric effects, calculations suggest that the position accuracy R_{xyz} of such a QPS can be below the 1 cm-level for an error (standard deviation) in the optical delays $\sigma_s = 1.0 \mu\text{m}$ associated with the minima in two-photon counting rates R_c . The complex dependence of R_{xyz} on user position suggests that significant engineering must be done to design a realistic QPS.

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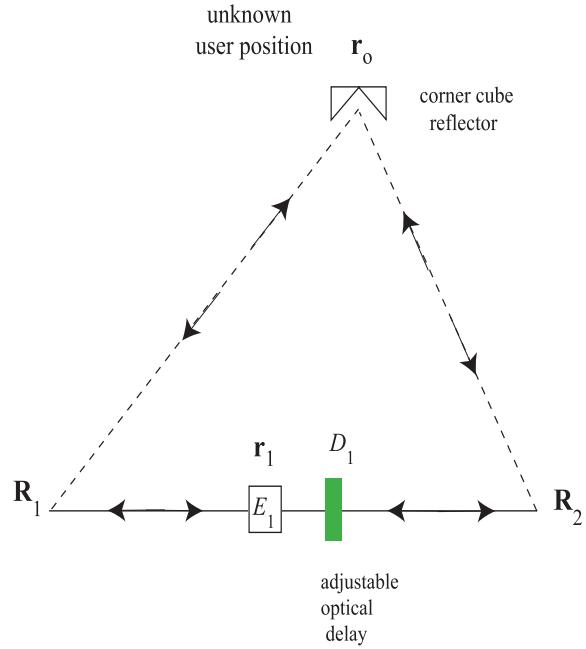


Figure 1: One baseline is shown for the quantum positioning system (QPS). Points \mathbf{R}_1 and \mathbf{R}_2 on the baseline contribute to the definition of the reference frame for spatial positioning. Box E_1 contains an entangled photon (biphoton) source and 50:50 beam splitter; see Figure 2. The quantity D_1 is a controllable, calibrated optical delay.

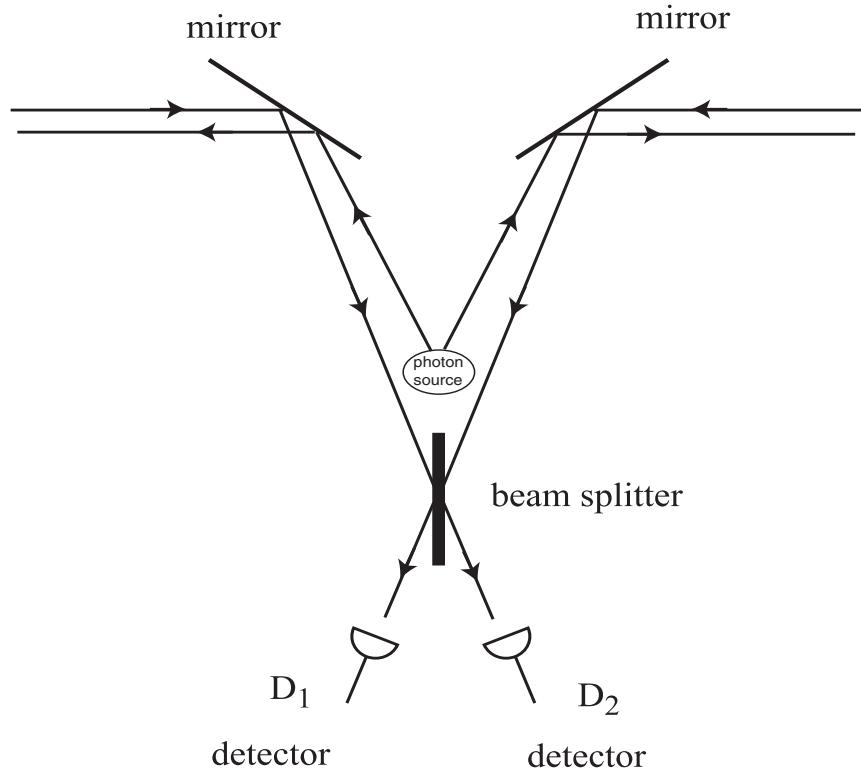


Figure 2: An expanded view of the contents of each of the three boxes E_1 , E_2 , and E_3 , which are located on the three baselines, one of which is shown in Figure 1. Each box contains an entangled photon (biphoton) source, a 50:50 beam splitter, and two single-photon detectors D_1 and D_2 , to perform photon coincidence counting.

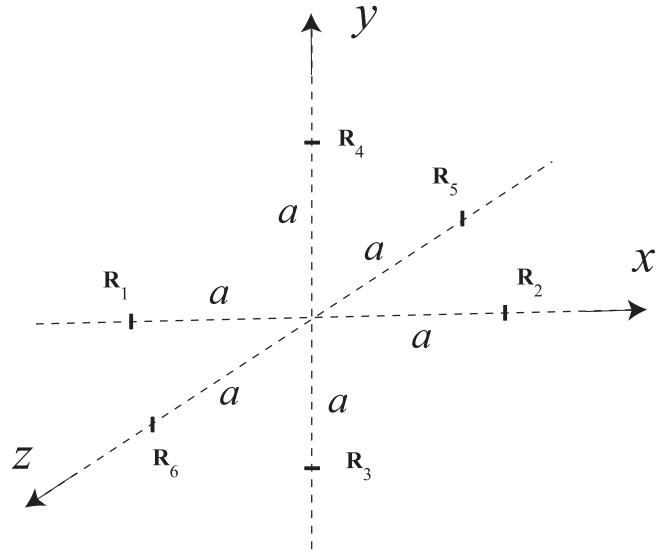


Figure 3: The baselines are shown for a possible terrestrial QPS that might be used on the Earth. The baselines lie along the x , y , and z axes, are of length $2a$ and are orthogonal to each other.

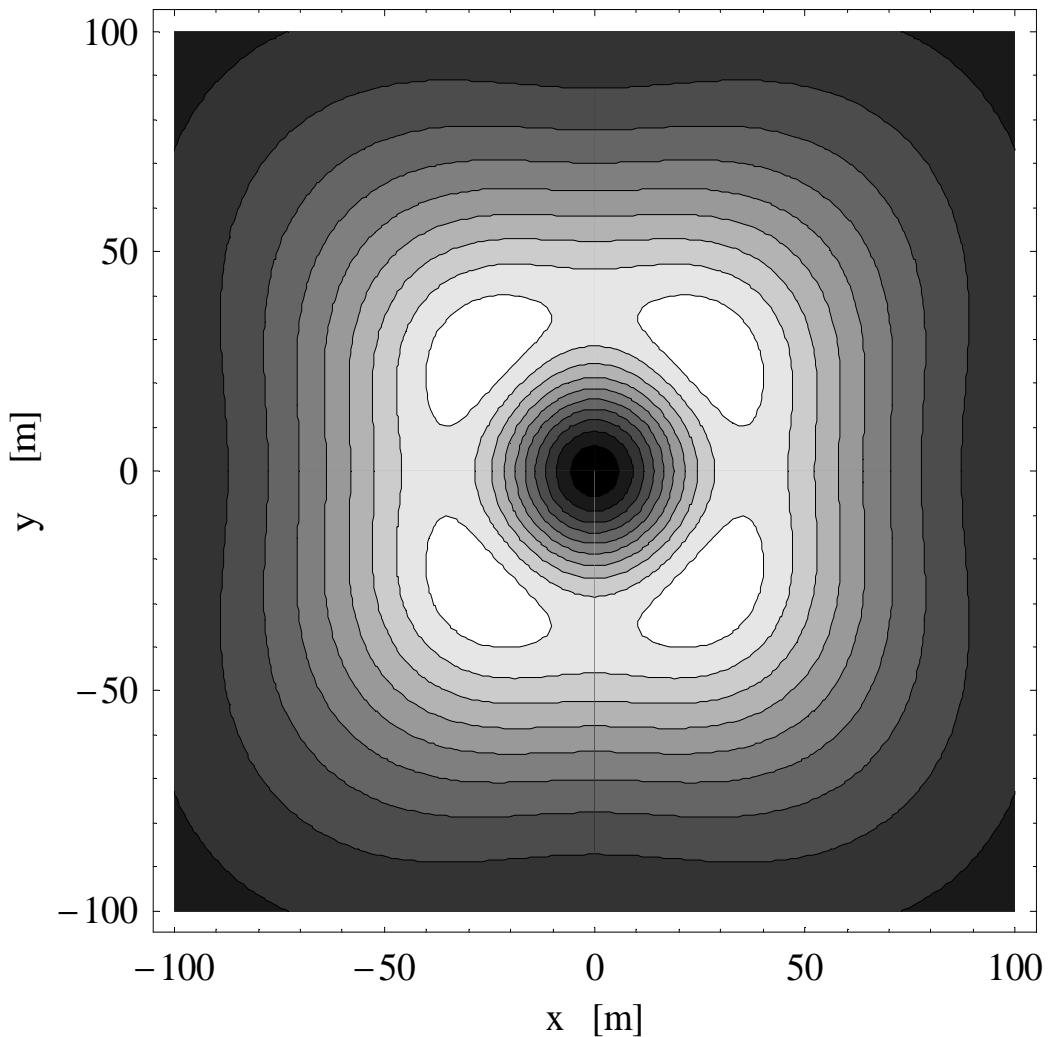


Figure 4: A plot of the contours of constant $1/R_{xyz}$ is shown in the $x_o - y_o$ plane at $z_o = 100/\sqrt{3}$ m. Light-shaded areas are small values of R_{xyz} . Units on both axes are meters.

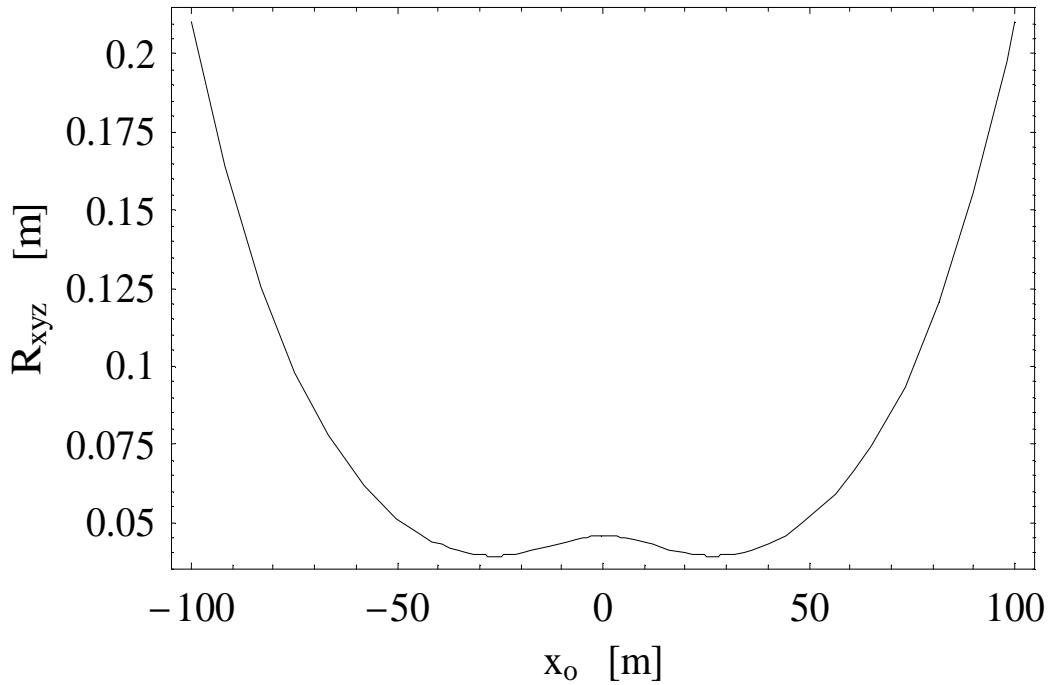


Figure 5: The error R_{xyz} vs. x_o is plotted for $y_o = 30$ m and $z_o = 100$ m/ $\sqrt{3}$, which corresponds to a line in Figure 4 with relatively small error R_{xyz} . The same parameters are used in this plot as in Figure 4. Units on both axes are meters.

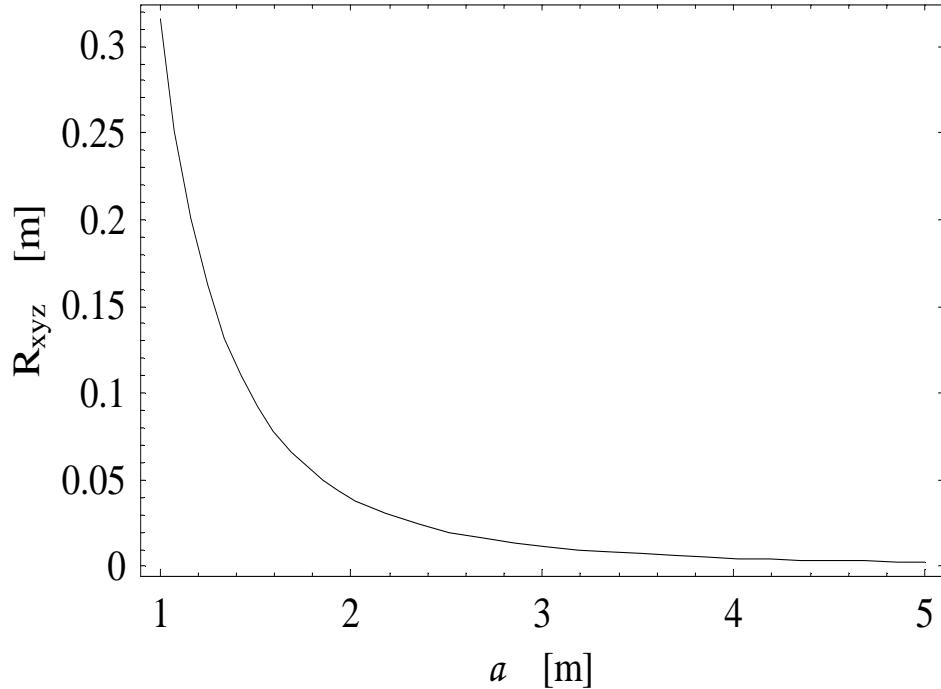


Figure 6: The error R_{xyz} vs. a (half the interferometer baseline length) is plotted for $x_o = y_o = 30$ m, and $z_o = 100$ m/ $\sqrt{3}$, which corresponds to the high-accuracy light-shaded region in upper right of Figure 4. Units on both axes are meters.

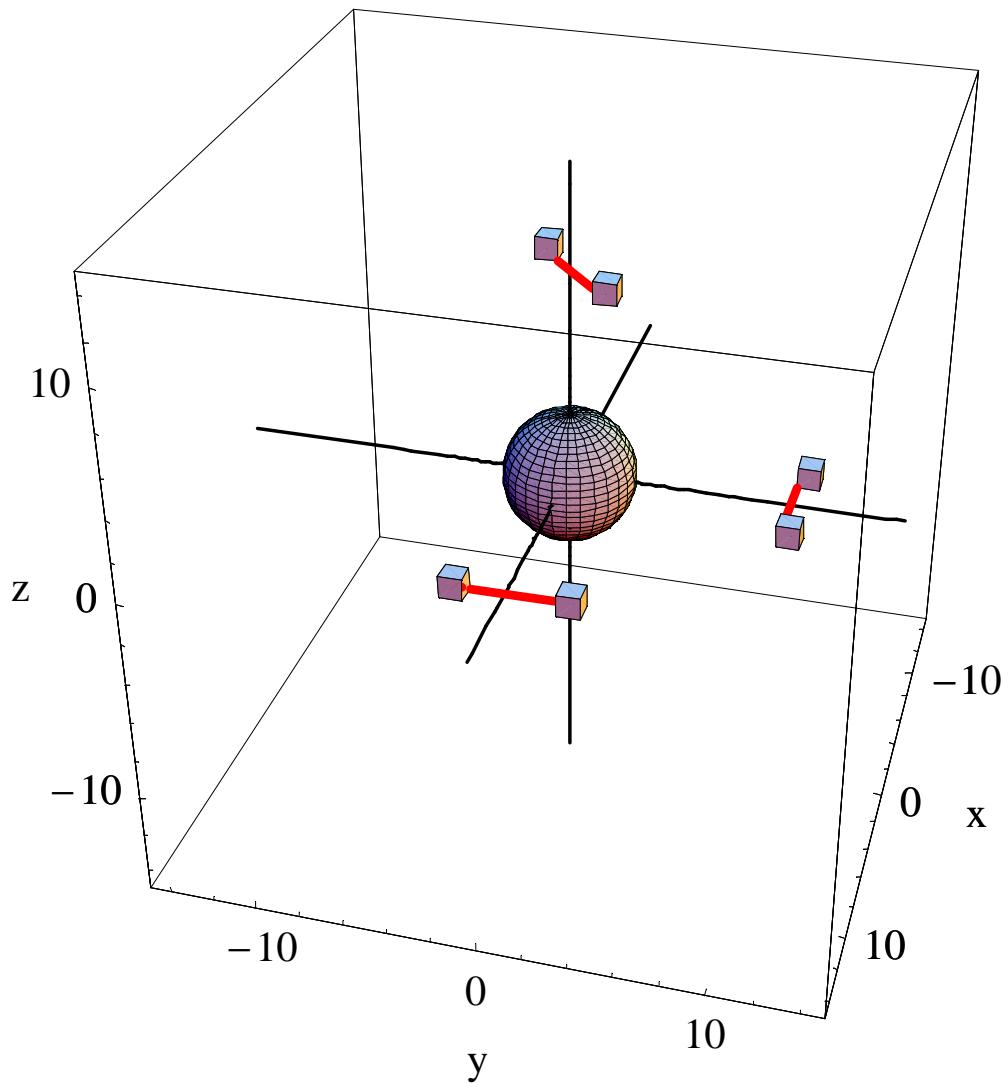


Figure 7: A schematic of the LEO satellite QPS is shown. Pairs of satellites orbiting Earth, shown by connecting lines, form the interferometer baselines of length b . Example numbers used in this calculation have baseline $b = 20$ km and LEO satellite semi-major axis $a = 7360$ km .

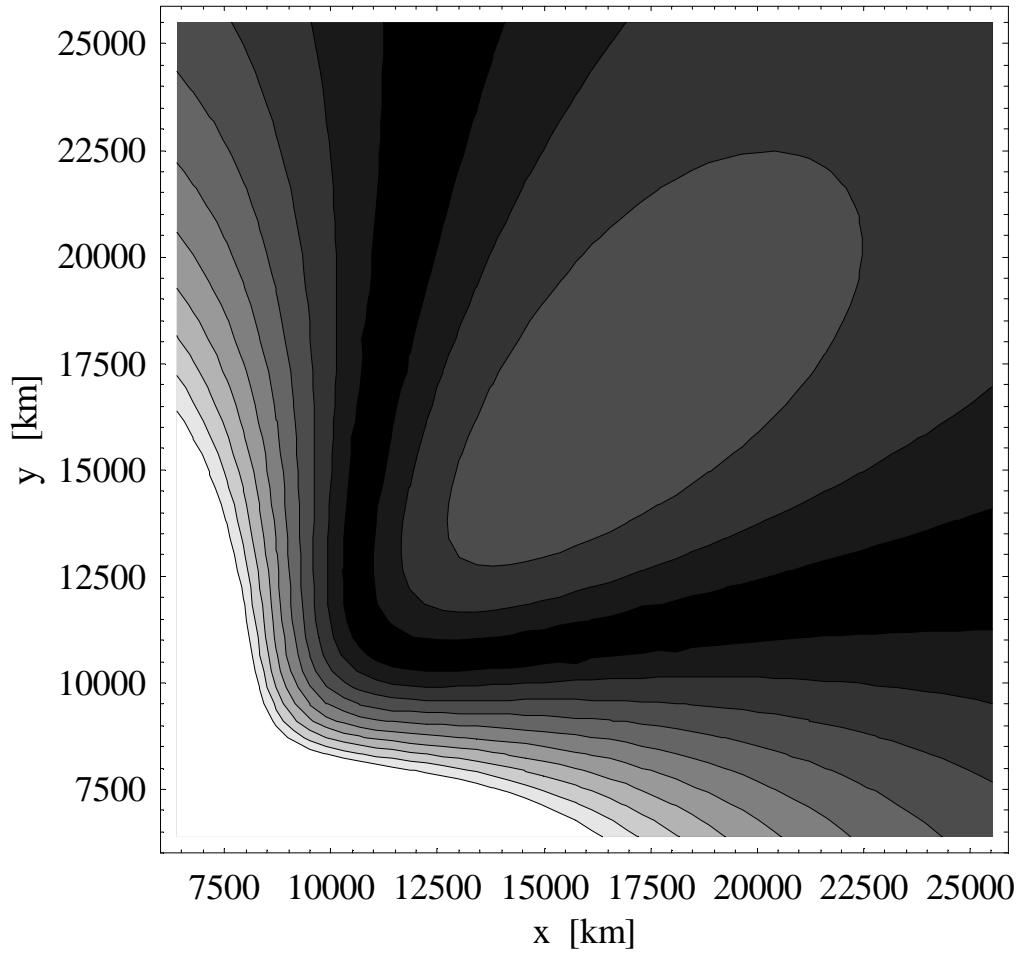


Figure 8: Contours of constant reciprocal position error, $1/R_{xyz}$, are shown in the $x_o - y_o$ plane for $z_o = R_e/\sqrt{3}$ and $\sigma_s = 1.0 \mu\text{m}$. Lighter-shaded areas are smaller values of error R_{xyz} . The semi-major axis of the LEO satellites is taken to be $a = 7360 \text{ km}$ and the baselines (satellite pair separation) $b = 20 \text{ km}$.

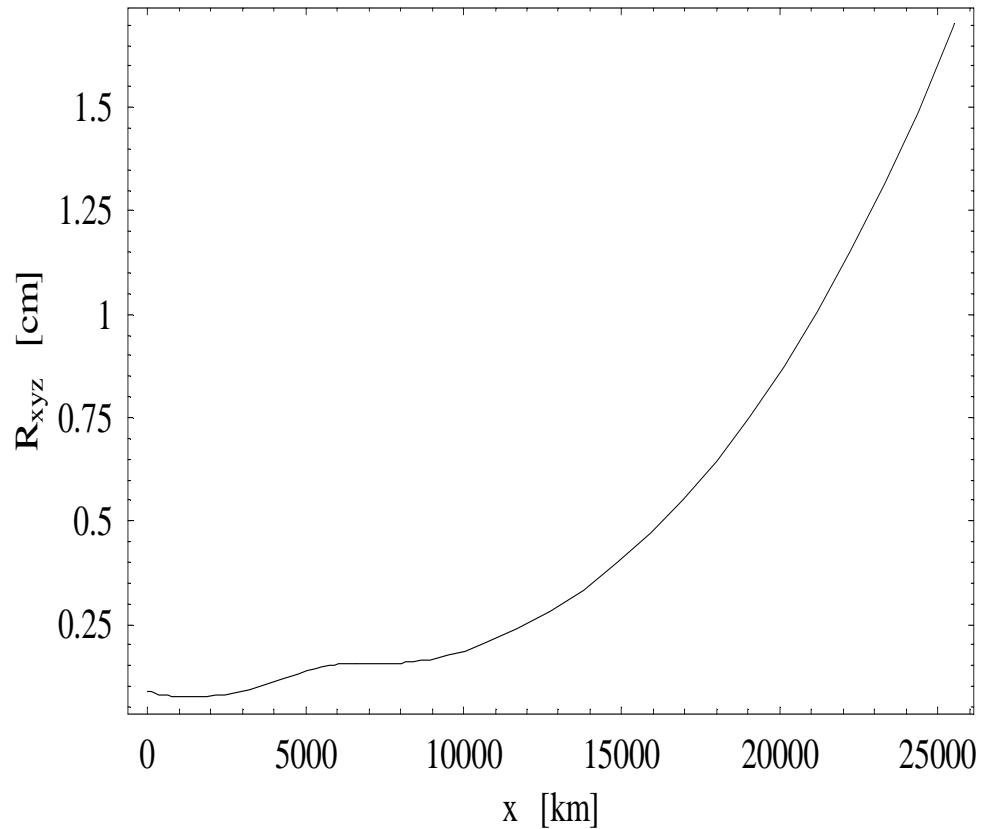


Figure 9: Plot of position error, R_{xyz} vs. x_o , shown for $y_o = z_o = R_e/\sqrt{3}$ with $\sigma_s = 1.0 \mu\text{m}$. The semi-major axis of the LEO satellites is taken to be $a = 7360 \text{ km}$ and the baselines (satellite pair separation) $b = 20 \text{ km}$.

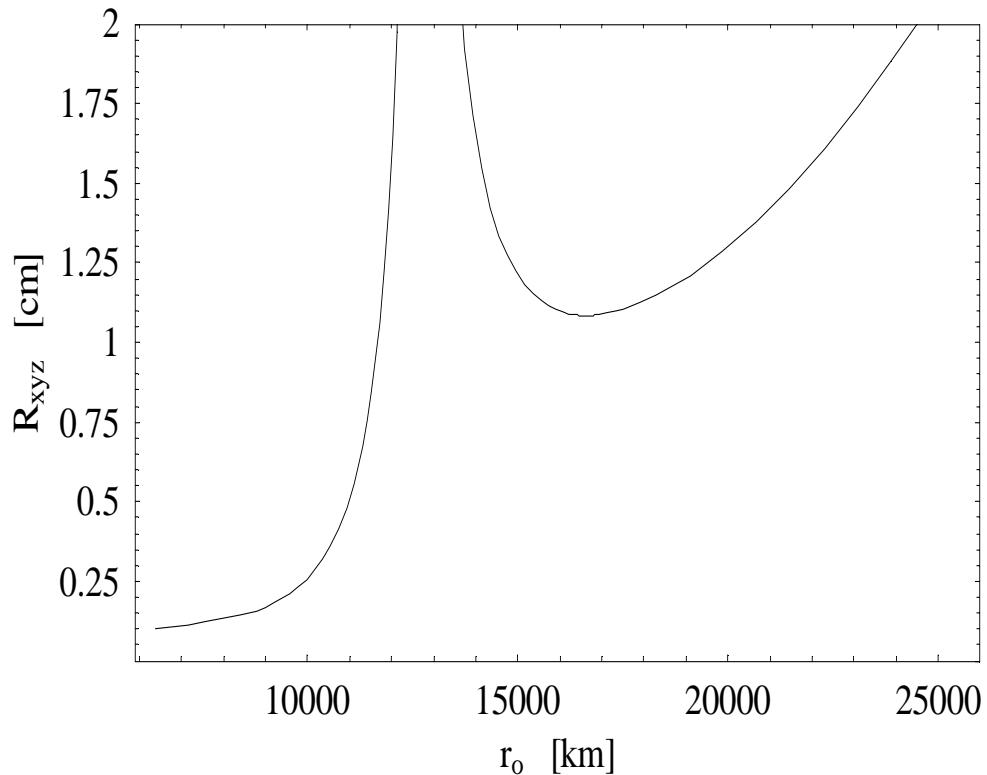


Figure 10: Plot of position error in the radial direction, R_{xyz} vs. r_o , where $r_o = \sqrt{x_o^2 + y_o^2 + z_o^2}$ and $\sigma_s = 1.0 \mu\text{m}$. The semi-major axis of the LEO satellites is taken to be $a = 7360 \text{ km}$ and the baselines (satellite pair separation) $b = 20 \text{ km}$.

QUESTIONS AND ANSWERS

JUDAH LEVINE (National Institute of Standards and Technology): I'm familiar with that paper, in fact, and as I understand that paper, it depends on the fact that the parameters of the medium — the propagation delay — is well understood and well controlled. And that the idea of extending it into the atmosphere, which is what those guys proposed, does not satisfy this implicit requirement.

THOMAS BAHDER: Right. In other words, if you tried to do Yanhua Shih's Scheme in vacuum, it would fail, because Yanhua Shih's scheme depends on knowing the exact group velocity in the two channels of the two fibers to the two clocks. So I didn't really describe that, I did not have time.

But this scheme, as you see, works in vacuum, right? I have to make a comment on that in that paper; it is described as if it is a one-way scheme. Well, if you look at relativity theory, you cannot synchronize clocks one way. It's logically inconsistent.

So really, in order to calibrate your system and measure the group velocity of each channel, you have to do a round-trip flight. But that could have been done at an earlier time if your fibers were stable. So this scheme would work in vacuum, presumably.

WLODZIMIERZ LEWANDOWSKI (Bureau International des Poids et Mesures): I have one question. You say you neglected most of those errors. Have you an idea about ...?

BAHDER: My funding has run out and so, unless I can get further funding for this, this is the end of the show.

LEWANDOWSKI: But it can be much larger than what you are showing here.

Yes. There are actually some very interesting things in that. The atmospheric effects can be huge. On the other hand, this single photon propagation had shown that you can cancel dispersion. You cannot cancel propagation delays, but the dispersion of the group velocity is cancelled out. It is a very weird effect.

So there can be some games that may be possible to play. And I am not sure ...

LEWANDOWSKI: And you need more money.

BAHDER: Yes, I need more money.